



Australian Government
Bureau of Meteorology

SPACE WEATHER SERVICES

SOLAR

AND

GEOPHYSICAL

SUMMARY

May 2018

Solar Activity

Solar activity was at very low levels this month, with no M or X class flares observed .

DATE	FLARES		FLARE MAX	FADEOUT POSSIBLE ON DAYLIGHT CIRCUIT
	CLASS M	CLASS X		
None observed.				

FLARE ALERT	SWF ALERT
None issued.	

PROTON ALERT						
DATE	BEGIN TIME	DATE	END TIME	ENERGY THRESHOLD		
None issued.						

Ionospheric Activity

Several broad periods of lower than normal MUFs was observed during May, particularly in the northern region of Australia, probably in association with lower solar flux levels.

Ionospheric depressions listed below are calculated using foF2 data from the SWS ionosonde network.

IONOSPHERIC DISTURBANCES (MUFs for the Australian region)

01 May 18 Northern region MUFs Depressed by 20%
02 May 18 Northern region MUFs Depressed by 15%
03 May 18 Northern region MUFs Depressed by 20%
04 May 18 Northern region MUFs Depressed by 15%
05 May 18 Northern region MUFs Depressed by 15%
06 May 18 Southern region MUFs Depressed by 15%
10 May 18 Southern region MUFs Depressed by 15%
11 May 18 Northern region MUFs Depressed by 15%
12 May 18 Southern region MUFs Depressed by 20%
13 May 18 Southern region MUFs Depressed by 15%
15 May 18 Northern region MUFs Depressed by 15%
19 May 18 Northern region MUFs Depressed by 15%
Southern region MUFs Depressed by 15%
20 May 18 Northern region MUFs Depressed by 15%
21 May 18 Northern region MUFs Depressed by 15%
22 May 18 Northern region MUFs Depressed by 15%
24 May 18 Southern region MUFs Depressed by 15%
30 May 18 Northern region MUFs Depressed by 15%

RADIO COMMUNICATIONS WARNINGS

DATE	Nº.	BEGIN	END
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06 May 18	06	06 May 18	08 May 18
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Geomagnetic Activity

The strongest disturbance for the month was observed on 06 May with a local A index of 19 reached. This activity was associated with the Earth's entry into a high speed solar wind stream induced by a coronal hole. This was the only significant geomagnetic disturbance for the month.

Disturbances with Australian A indices greater than or equal to 16 are reported below.

DATE	GEOMAGNETIC DISTURBANCES (for the Australian region)
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06 May 18	Quiet to Active
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GEOMAGNETIC WARNINGS AND ALERTS				
DATE	N°.	BEGIN	END	ISSUED

06 May 18	09	06 May 18	08 May 18	Warning
15 May 18	10	16 May 18	17 May 18	Warning
30 May 18	11	01 Jun 18	01 Jun 18	Warning

Solar And Geophysical Indices

DATE	10cm FLUX	AUSTRALIAN A-INDEX	AUSTRALIAN T-INDEX	DATE	10 cm FLUX	AUSTRALIAN A-INDEX	AUSTRALIAN T-INDEX
01 May	68	1	1	17 May	69	6	7
02 May	67	1	-3	18 May	69	3	9
03 May	67	2	-2	19 May	70	0	6
04 May	68	1	3	20 May	69	1	2
05 May	68	10	1	21 May	70	0	1
06 May	67	19	20	22 May	71	3	5
07 May	70	11	12	23 May	73	5	11
08 May	70	12	9	24 May	74	2	4
09 May	70	10	3	25 May	76	1	4
10 May	70	9	5	26 May	73	0	2
11 May	70	9	5	27 May	75	2	3
12 May	70	5	4	28 May	77	1	4
13 May	71	6	1	29 May	75	1	9
14 May	70	2	-2	30 May	75	1	5
15 May	70	1	-1	31 May	77	7	4
16 May	70	1	4				

DATE	10cm FLUX	AUSTRALIAN A-INDEX	AUSTRALIAN T-INDEX	SUNSPOT NUMBER	FLARES	
	Monthly Average	Monthly Average	Monthly Average	Monthly Average	13-month smoothed	>M1.0
Jun 17	74.8	3.4	9.2	19.4	22.3	0
Jul 17	78	5.8	10.5	18.3	21	3
Aug 17	78	7.3	10.4	33.1	19.5	1
Sep 17	93.4	10.4	19.5	43.6	18.2	31
Oct 17	76.4	6.9	4.6	13.2	16.6	1
Nov 17	72.2	7.1	-5.6	5.7	15.4	0
Dec 17	71.7	5.8	-9.2	8.2	14.8	0
Jan 18	70	4.8	-6.5	6.7	13.9	0
Feb 18	72.1	5	-0.9	10.6	12.5	0
Mar 18	68.4	5.4	2.4	2.5	9.9	0
Apr 18	70.1	4.4	6.5	8.9	8.1	0
May 18	70.9	4.3	4.4	13.2	7.8	0

Predicted sunspot numbers

SPECIAL NOTE

In June 2015, WDC-SILSO, Royal Observatory of Belgium, Brussels, adjusted their original observed sunspot numbers (Version 1.0) to a Version 2.0 series in which the sunspot numbers are higher. In the Version 1.0 data series, some of the sunspot numbers were weighted. The new (Version 2.0) sunspot numbers are unweighted sunspot number counts (for more information see: www.sidc.be/silso/home). The Bureau of Meteorology SWS solar cycle prediction is now based on the Version 2.0 values. As a result of this recalibration, the observed monthly and observed and predicted smoothed sunspot numbers are higher in the monthly table and solar cycle graph.

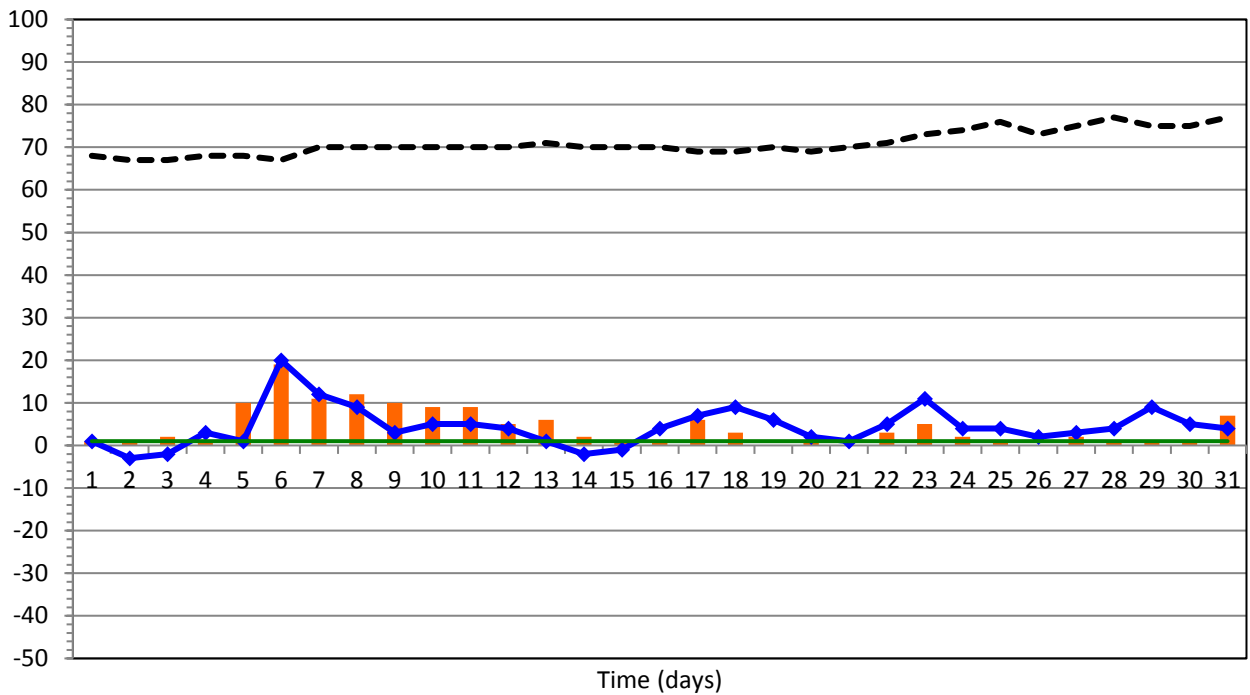
SWS WORLD T-INDICES

Year	Jan	Feb	Mar	Apr	May	Jun	Jul	Aug	Sep	Oct	Nov	Dec
2013	74	62	69	84	92	82	81	76	68	76	86	90
2014	91	108	130	114	96	84	86	81	90	94	98	103
2015	92	101	91	97	85	79	71	48	40	43	53	50
2016	52	57	46	30	34	28	26	25	28	17	12	10
2017	14	19	11	8	2	2	0	3	16	13	5	6
2018	8	3	3	-2	-7*	-1	-2	-3	-3	-4	-4	-5
2019	-5	-5	-5	-4	-4	-3	-3	-2	-1	0	1	3
2020	5	7	9	11	14	17	20	23	26	28	32	35
2021	38	40	42	44	46	49	50	51	52	54	55	57

SWS predicted T-index

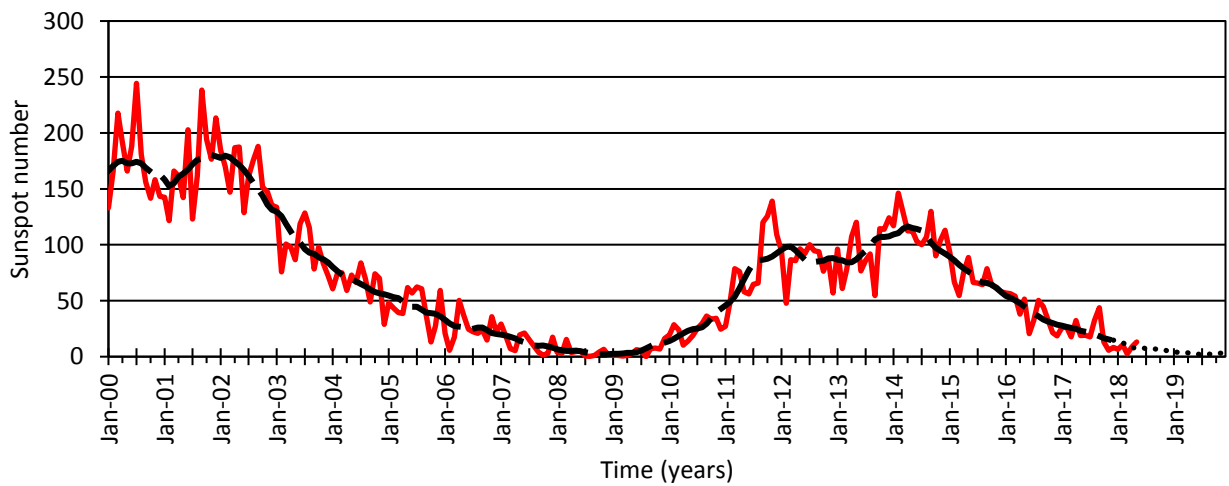
* New observed data. T-index value may have changed

Solar and Geophysical Indices - May 2018



■ Australian A-index 10cm flux
◆ Australian observed T index Global predicted monthly T index

Solar Cycle



— Observed monthly sunspot number
 Smoothed monthly sunspot number
 Predicted smoothed monthly sunspot number

Observed monthly sunspot numbers courtesy of WDC-SILSO, Royal Observatory of Belgium, Brussels (www.sidc.be/silso/home)

NOTES - notes - NOTES - notes - NOTES - notes

- a. Times quoted in this publication are all Universal Time (UT).
- b. The values of all indices are provisional. Final values are not available for several months.
- c. M or X class flares refer to the X-ray classification system for solar flares. In this system, X class flares are more energetic than M class flares.

FLARE CLASS	X-RAY FLUX DENSITY	
	Ergs/cm ² /sec	W/m ²
M	0.01-0.1	10 ⁻⁵ to 10 ⁻⁴
X	>0.1	> 10 ⁻⁴

- d. Class M flares, particularly the less energetic ones, are likely to cause a fadeout on only the lowest frequencies. Class X flares are likely to cause a fadeout over the entire HF spectrum. It should be noted that a fadeout will only occur on those circuits having a reflection point in the daylight hemisphere of the earth. Circuits having only night hemisphere reflection points will not be affected no matter the energy of the solar flares.
- e. The 10cm flux is the radio power of the sun at a frequency of 2800 MHz (wavelength 10.7 cm). This flux is a good indicator of solar activity and is widely used in place of the sunspot number. The values are measured by the Penticton radio observatory, Canada. Unlike the sunspot number, the 10cm flux never drops to zero even during solar minimum. With no sunspots visible on the solar disk, the 10cm flux will still have a value of around 67. The table below gives a (statistical) comparison between 10cm flux and sunspot number. The 10cm flux is measured in solar flux units (10⁻²² W m⁻² Hz⁻¹).

SUNSPOT No.	10 cm FLUX
0	67
20	78
40	93
60	110
100	147
150	195
200	243

- f. Ionospheric disturbances refer to measurements made across Australia, but are generally applicable to mid-latitude Southern Hemisphere conditions. Spread F conditions indicate tilts in the ionosphere,

which may result in multipath fading on some HF circuits.

- g. The magnetic A-indices are for the Australian region. Large values for the A index correspond to disturbed conditions. Levels of magnetic disturbances are described in the following terms.

A INDEX VALUE	DESCRIPTION
0 up to 7	Quiet
8 up to 15	Unsettled
16 up to 24	Active
25 up to 35	Minor Storm
36 and above	Major Storm

- h. The Australian daily T-index is a measure of the average of the ionospheric critical frequencies available on a particular day - the higher the value of the T-index, the higher the ionospheric critical frequencies (and Maximum Usable Frequencies on HF circuits) for that day. The T-index is based on data from Australian ionospheric stations and so is most applicable to HF circuits with reflection points in the Australian region.
- i. The SWS monthly observed T-index is derived from the observed monthly median values of foF2 for each hour from ionospheric stations worldwide.

The predicted smoothed monthly T-indices are computed by using a statistical analysis of the observed monthly T-indices for all solar cycles since 1938.

The SWS T-indices may not be updated each month but only when sufficient new data becomes available.

SPACE WEATHER SERVICES
 PO BOX 1386, HAYMARKET NSW 1240
 A U S T R A L I A

GENERAL ENQUIRIES: +61 2 9213 8000

DUTY FORECASTER: +61 2 9213 8010

FACSIMILE: +61 2 9213 8060

E-mail: asfc@bom.gov.au

Web: www.sws.bom.gov.au